

1. 40 points.

- (a) Which statements are true for a perfect blackbody? **Solution:** As T increases, more photons are radiated (and so more energy) at all frequencies, but the maximum shifts towards higher frequency. As T increases:

- i. more energy is radiated at low frequency. T
- ii. more energy is radiated at high frequency. T
- iii. the most intense radiation shifts to longer wavelength. F It shifts to higher *frequency*.
- iv. the photon number *and* average energy both increase. T

- (b) Add the periodic boundary conditions and calculate the energy of the Ising model lattice below.

$$\begin{array}{ccccccc} + & - & - & & 6 & -4 & -4 \\ - & + & - & \Rightarrow & -4 & 6 & -4 \\ - & - & - & & -4 & -4 & -4 \end{array} = \frac{-16D}{2} = \boxed{-8D}.$$

- (c) Two gas-phase samples, Ar ($\sigma = 36 \text{ \AA}^2$, $m = 40 \text{ amu}$) and C_2H_2 ($\sigma = 72 \text{ \AA}^2$, $m = 26 \text{ amu}$), have the same pressure and temperature. Given the values for Ar below, write the corresponding values for C_2H_2 in the same units.

parameter	Ar	C_2H_2	
$\rho (\text{m}^{-3})$	$4.3 \cdot 10^{23}$	$4.3 \cdot 10^{23}$	depends only on P and T
$\lambda (\text{m})$	$6.5 \cdot 10^{-6}$	$3.2 \cdot 10^{-6}$	inversely prop. to σ
$\langle v_{\text{AA}} \rangle (\text{m s}^{-1})$	696	863	prop. to $\sqrt{1/m}$
$\gamma (\text{s}^{-1})$	$1.1 \cdot 10^8$	$2.7 \cdot 10^8$	equal to $\rho \langle v_{\text{AA}} \rangle \sigma$

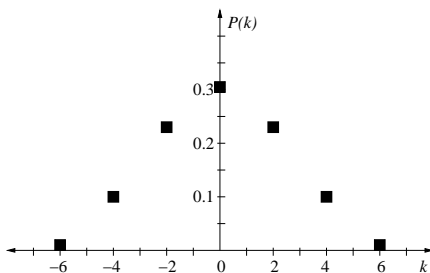
- (d) The diffusion constant for SF_6 in air is $0.150 \text{ cm}^2 \text{ s}^{-1}$ at 373 K. Find r_{rms} after 1 hour.
Solution: Use the Einstein diffusion equation:

$$\langle r^2 \rangle^{1/2} = \sqrt{6Dt} = [6(0.150 \text{ cm}^2 \text{ s}^{-1})(3600 \text{ s})]^{1/2} = \boxed{57 \text{ cm}}.$$

2. Graph $\mathcal{P}(k)$ for **six** coin flips. **Solution:** Use the equation for the discrete distribution $\mathcal{P}(k)$ when N is small. Because $N = 6$ is even, only even values of k will be possible:

$$\begin{aligned} \mathcal{P}(k = \pm 6) &= \frac{6!}{2^6 6! 0!} = \frac{1}{64} = 0.0156 & \mathcal{P}(k = \pm 4) &= \frac{6!}{2^6 5! 1!} = \frac{6}{64} = 0.0937 \\ \mathcal{P}(k = \pm 2) &= \frac{6!}{2^6 4! 2!} = \frac{15}{64} = 0.2344 & \mathcal{P}(k = 0) &= \frac{6!}{2^6 3! 3!} = \frac{20}{64} = 0.3125 \end{aligned}$$

As a check, these probabilities for all 7 possible values of k add up to 1.



3. In an ideal gas sample, PV is one measure of the energy content of the sample. If we set $PV_0 = \langle E_{AB} \rangle$, where $\langle E_{AB} \rangle$ is the average energy per collision, find an expression for the characteristic volume V_0 in terms of the mean free path and collision cross section. **Solution:** This is a mix-and-match algebra problem. In terms of our collision parameters, P is related to ρ which in turn determines λ , and $\langle E_{AB} \rangle$ is equal to $3k_B T/2$. We can start from that and see where it leads us, keeping in mind that the final expression should depend on λ and σ only:

$$\begin{aligned} V_0 &= \frac{\langle E_{AB} \rangle}{P} = \frac{3k_B T/2}{\rho R T / \mathcal{N}_A} = \frac{3k_B T/2}{\rho k_B T} \\ &= \frac{3}{2\rho} = \boxed{\frac{3}{2}\lambda\sigma}. \end{aligned}$$

4. Find an equation for the blackbody $\rho(\nu)$ assuming the energy is stored in quantum **rotations**. **Solution:** Starting from the general Eq. 15.25,

$$\rho(\nu) = \frac{8\pi\nu^2 \langle \epsilon \rangle}{c^3},$$

we need the average rotational energy, rather than the average vibrational energy. To get that, we use the average value theorem and the partition function for rotation:

$$\begin{aligned} \langle \epsilon_{\text{rot}} \rangle &= \int_0^\infty \mathcal{P}(J) \epsilon_{\text{rot}} dJ = \frac{1}{q_{\text{rot}}} \int_0^\infty (2J+1) e^{-BJ(J+1)/(k_B T)} BJ(J+1) dJ \\ &= \frac{B}{k_B T} \int_0^\infty e^{-Bx/(k_B T)} Bx dx = \frac{B^2}{k_B T} \left(\frac{1}{[B/(k_B T)]^2} \right) \quad \text{set } x = J(J+1) \\ &= \frac{B^2}{k_B T} \left(\frac{k_B^2 T^2}{B^2} \right) = k_B T \end{aligned}$$

which we also could have predicted from the equipartition principle. Now we put this into the equation above for $\rho(\nu)$:

$$\rho(\nu) = \frac{8\pi k_B T \nu^2}{c^3}.$$

This diverges to infinity if you integrate over all frequencies, because the integral approximation to q_{rot} is not exact.